

Thermalization and Plasma Instabilities

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Strong and Electroweak Matter 2006

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Motivation

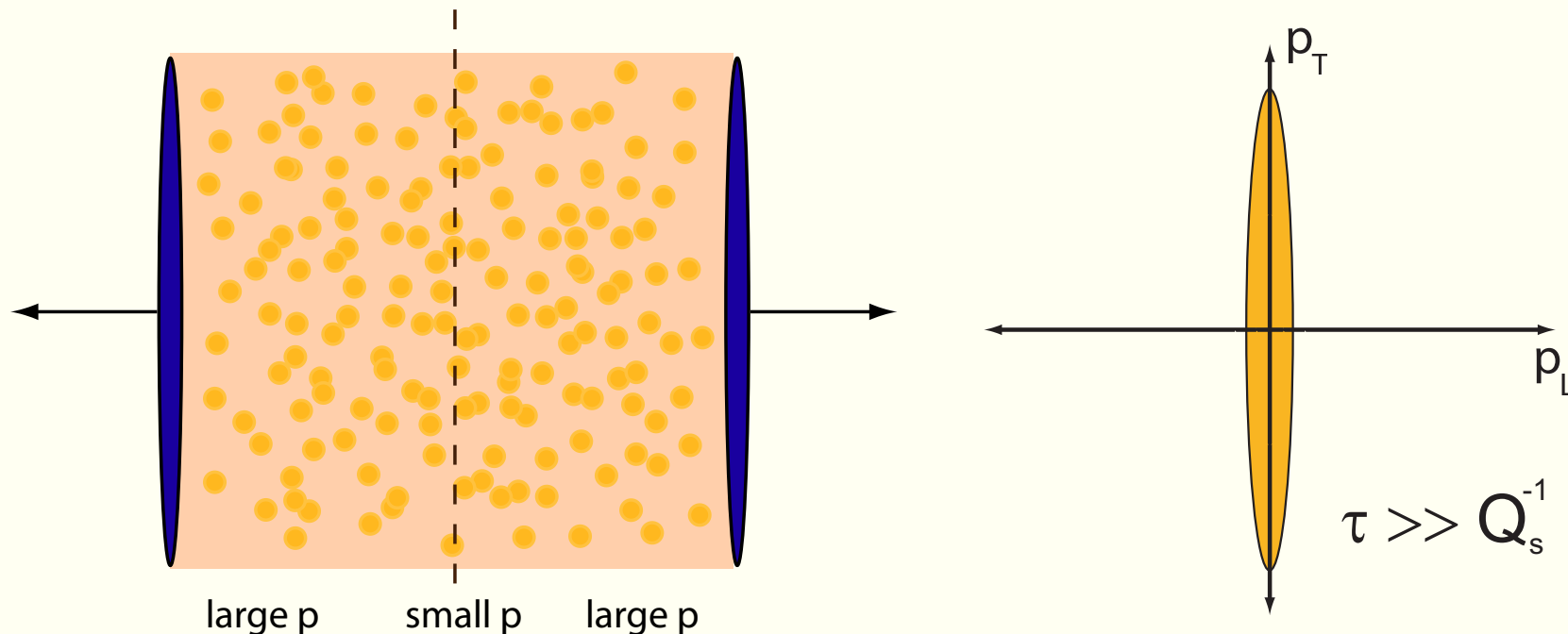
- We would like to have a first principles derivation of the mechanisms and time scales necessary for the isotropization and equilibration of a quark-gluon plasma.
- In addition to equilibration via 2-2 elastic scattering (super slow) one needs to include inelastic processes, e.g. bremsstrahlung 2-3 (and 3-2) processes, and the effect of background fields.
- In equilibrium the background field (soft modes) only serves to screen the interaction (Debye screening). However, in a non-equilibrium setting the background field can have non-trivial dynamics.
- Consider, for example, a spatially homogeneous plasma which has been initialized such that it has a “temperature” anisotropy.
- In such an anisotropic plasmas new collective modes corresponding to *electro-/chromodynamic instabilities* appear.

Why anisotropic distribution functions?

Because of the natural expansion of the system the gluon distribution functions created during relativistic heavy ion collisions are *generically* locally anisotropic in momentum space.

$$\langle p_T \rangle \sim Q_s \quad (\text{nuclear saturation scale})$$

$$\langle p_L \rangle \sim 1/\tau \quad (\text{free streaming})$$



Current Filamentation

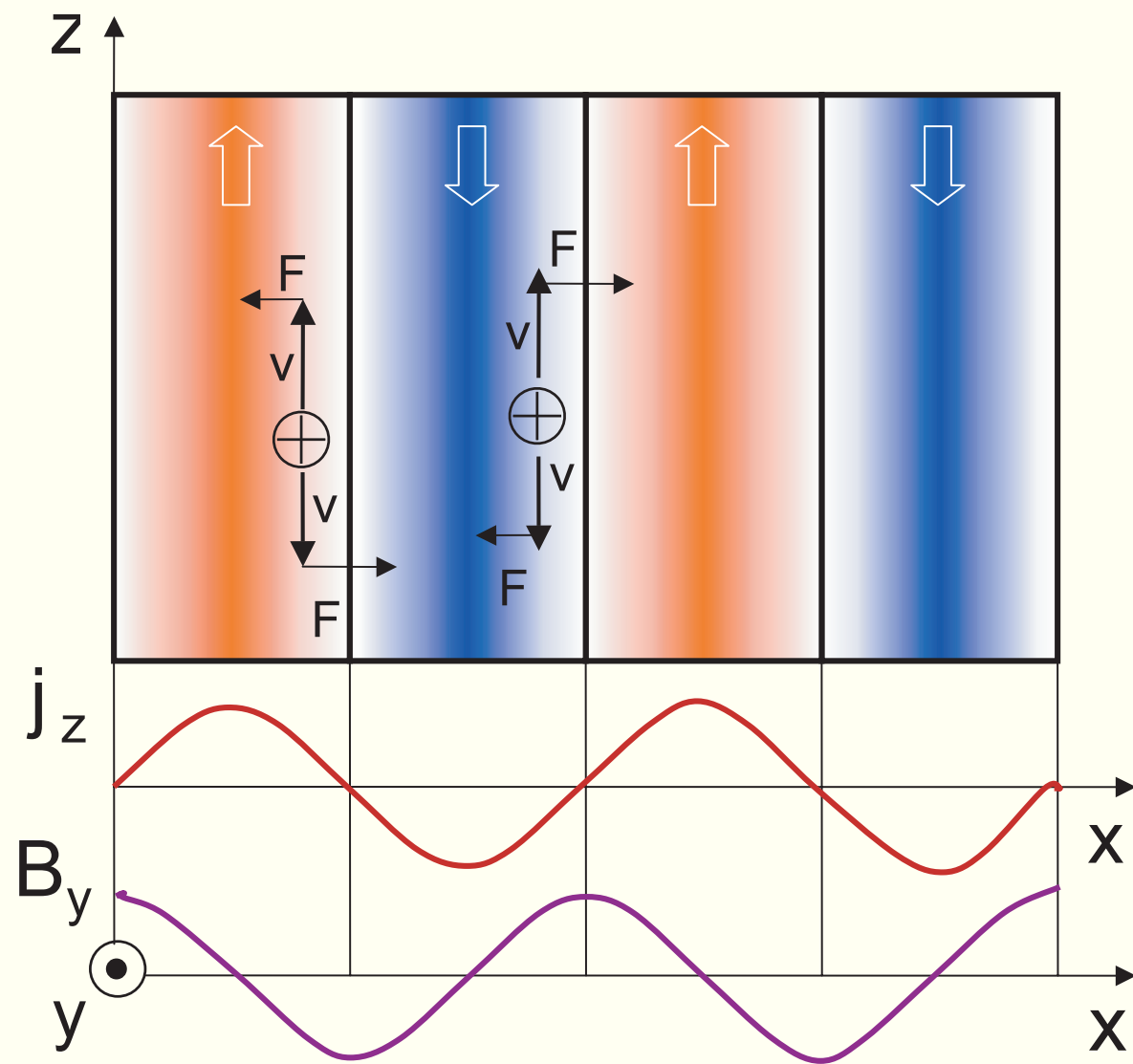
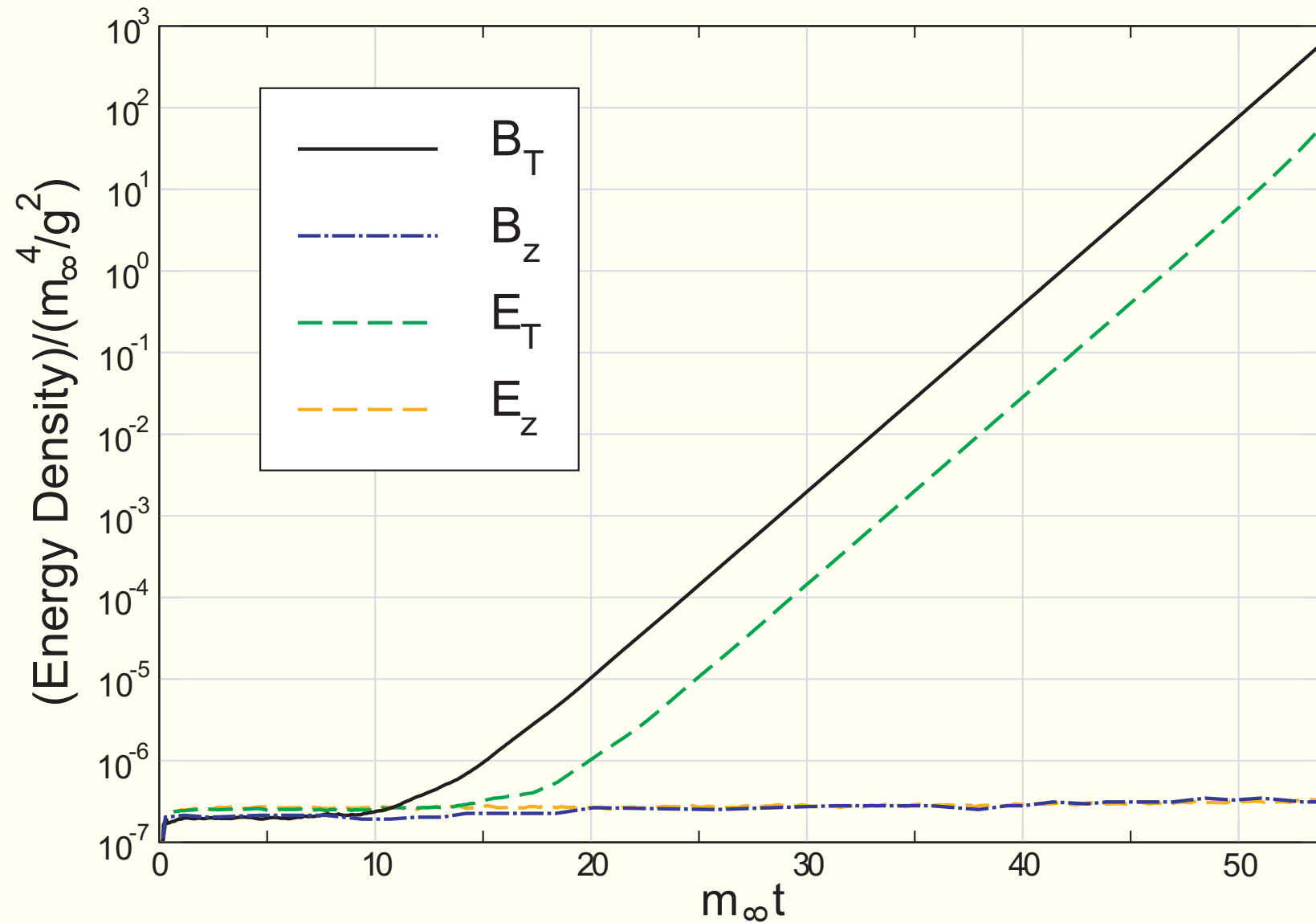


Figure courtesy S. Mrówczyński

Abelian Plasma



Collective Modes of an Isotropic QGP

The isotropic hard-thermal-loop (HTL) gluon propagator is given by

$$\Delta^{ij} = (k^2 - \omega^2 + \Pi_T)^{-1}(\delta_{ij} - k^i k^j / k^2) - \frac{k^2}{\omega^2} (k^2 - \Pi_L)^{-1} k^i k^j / k^2$$

with

$$\Pi_T(\omega, k) = \frac{m_D^2}{2} \frac{\omega^2}{k^2} \left[1 - \frac{\omega^2 - k^2}{2\omega k} \log \frac{\omega + k}{\omega - k} \right],$$

$$\Pi_L(\omega, k) = m_D^2 \left[\frac{\omega}{2k} \log \frac{\omega + k}{\omega - k} - 1 \right],$$

and $m_D \propto gT$.

$$\lim_{\omega \rightarrow 0} \Pi_L(\omega, k) = m_D^2 \quad \text{electric screening}$$

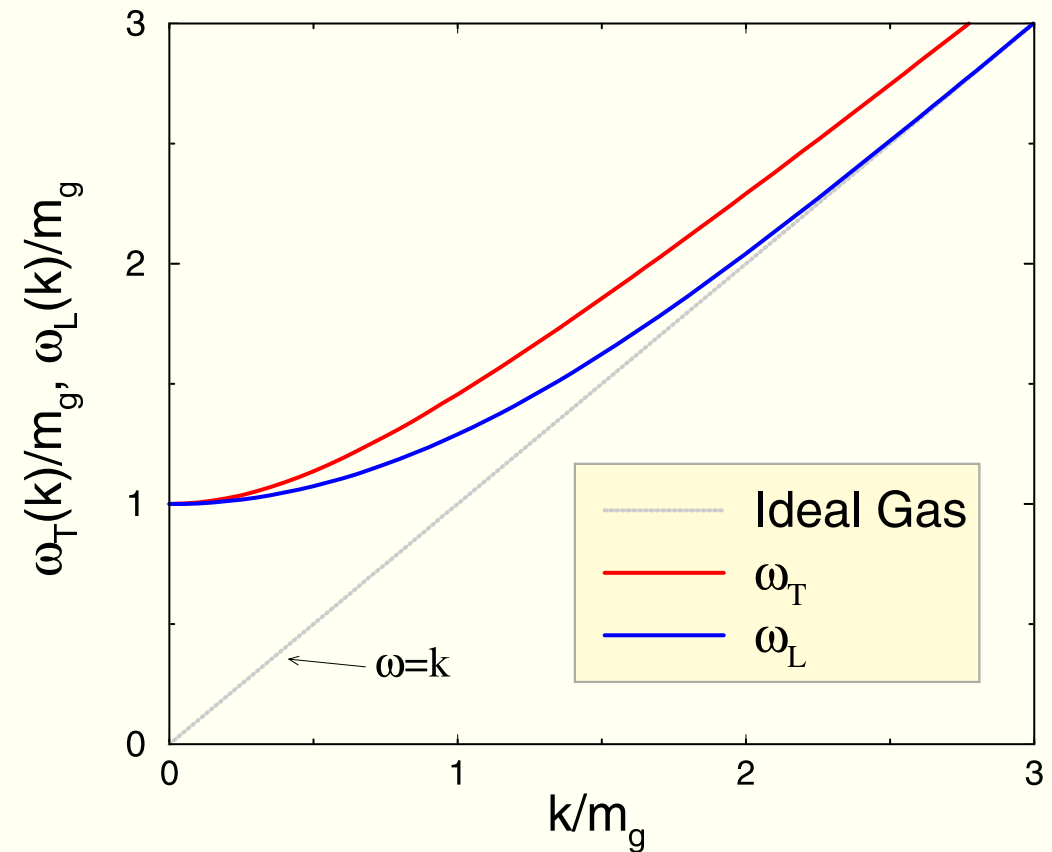
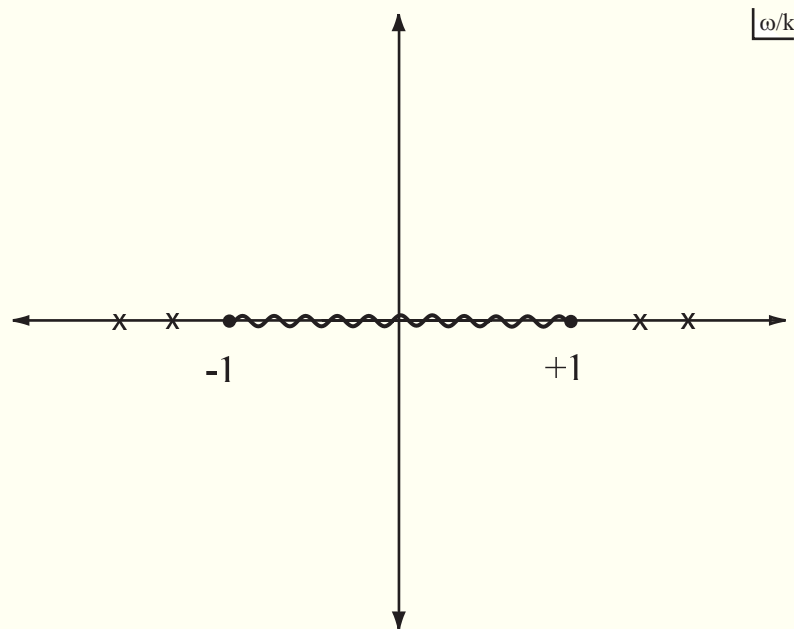
$$\lim_{\omega \rightarrow 0} \Pi_T(\omega, k) = 0 \quad \text{no magnetic screening}$$

Collective Modes of an Isotropic QGP

In the isotropic case the only poles are at real timelike ($\omega > k$) momentum. In order to determine the dispersion relations for these excitations we can then explicitly look for the poles in the propagator.

$$0 = k^2 - \omega_T^2 + \Pi_T(\omega_T, k)$$

$$0 = k^2 - \Pi_L(\omega_L, k)$$



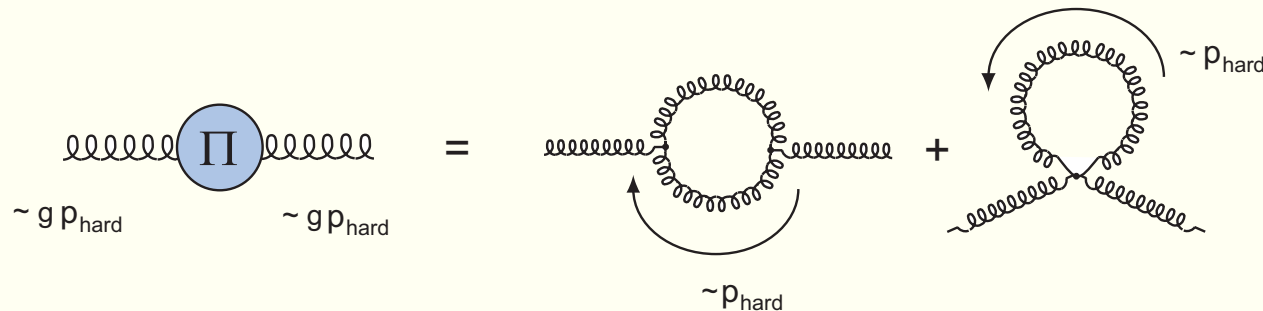
Anisotropic Gluon Polarization Tensor

In order to determine the HL gluon polarization we can use either linearized three-dimensional kinetic theory (Boltzmann-Vlasov eq)

$$[v \cdot D_X, \delta n(p, X)] + g v_\mu F^{\mu\nu}(X) \partial_\nu^{(p)} n(\mathbf{p}) = 0$$

$$D_\mu F^{\mu\nu} = J^\nu = g \int_p v^\nu \delta n(p, X)$$

or diagrammatically



In both cases the result for the retarded self-energy is

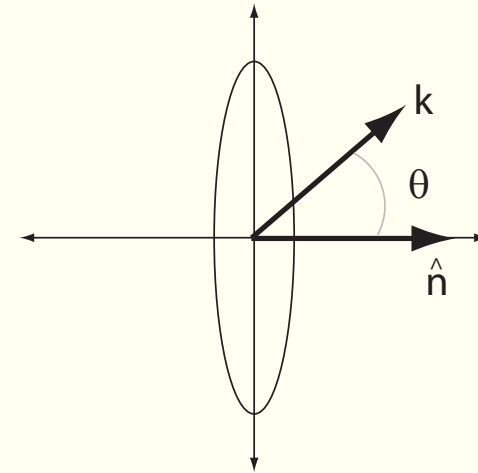
$$\Pi_{ab}^{ij}(K) = -g^2 \delta_{ab} \int_{\mathbf{p}} v^i \partial_l f(\mathbf{p}) \left(\delta^{jl} - \frac{v^j k^l}{K \cdot V + i\epsilon} \right)$$

The nature of the anisotropy

We assume that the anisotropic distribution function can be obtained from an arbitrary isotropic distribution function by a change of its argument.

$$f(p^2) \rightarrow f(p^2 + \xi(p \cdot n)^2)$$

The polarization tensor can then be written as

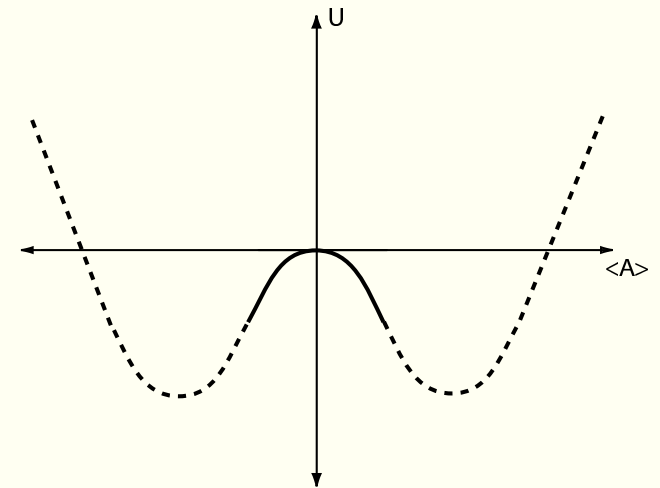
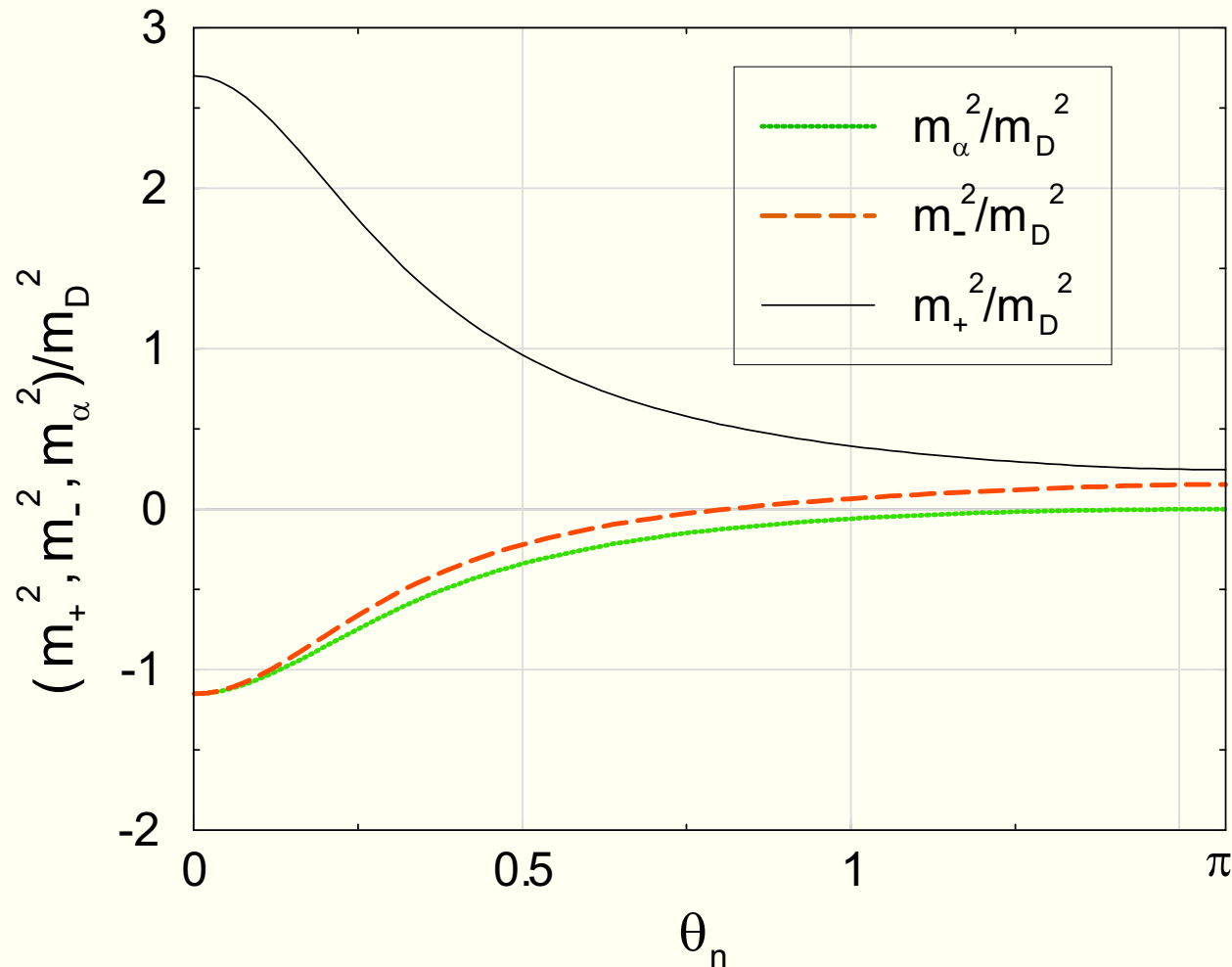


$$\Pi^{ij}(K) = m_D^2 \int \frac{d\Omega}{4\pi} v^i \frac{v^j + \xi(v \cdot n)n^j}{(1 + \xi(v \cdot n)^2)^2} \left(\delta^{jl} - \frac{v^j k^l}{K \cdot V + i\epsilon} \right)$$

where m_D is the *isotropic* Debye mass

$$m_D^2 = -\frac{g^2}{2\pi^2} \int_0^\infty dp p^2 \frac{df(p^2)}{dp}$$

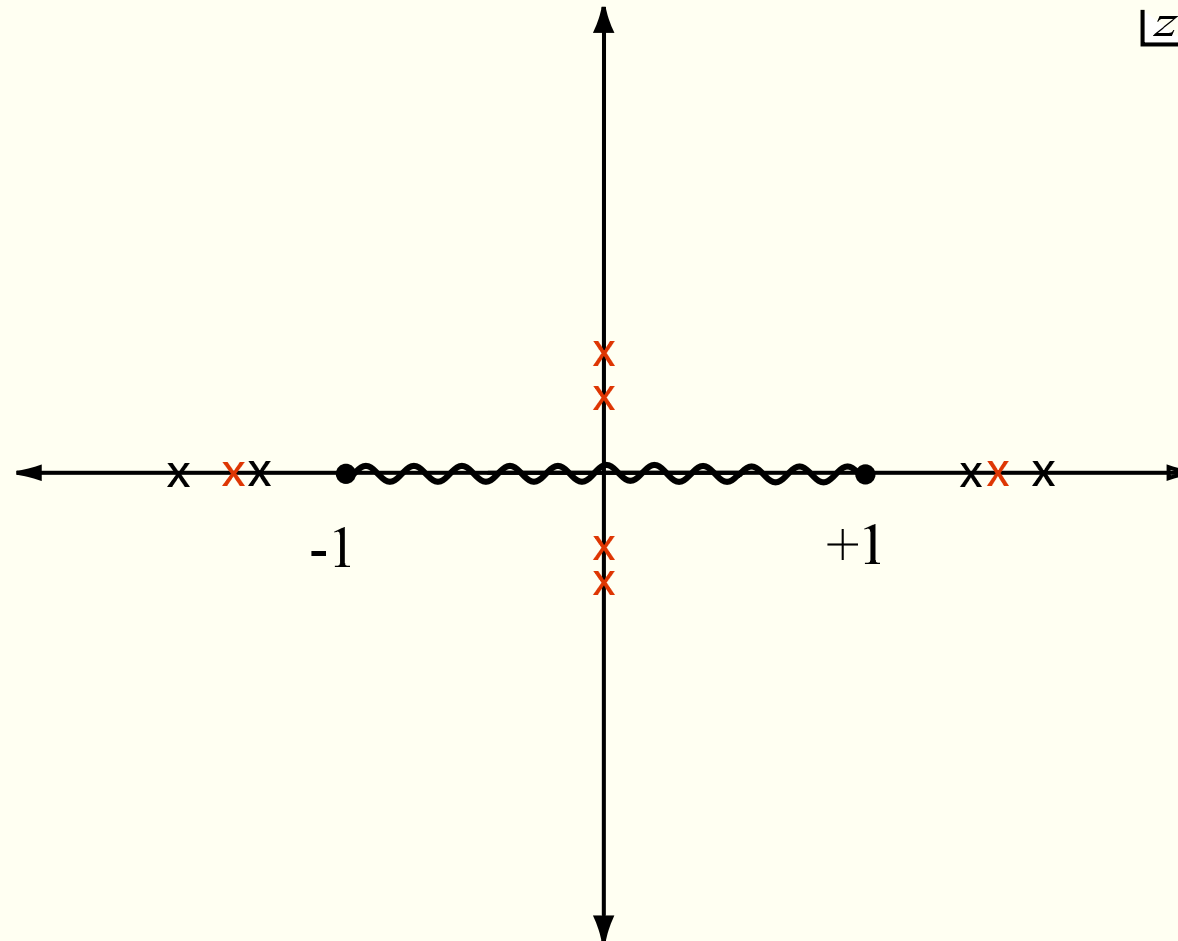
New Mass Scales – $\xi > 0$



Sketch of the effective potential of an unstable mode.

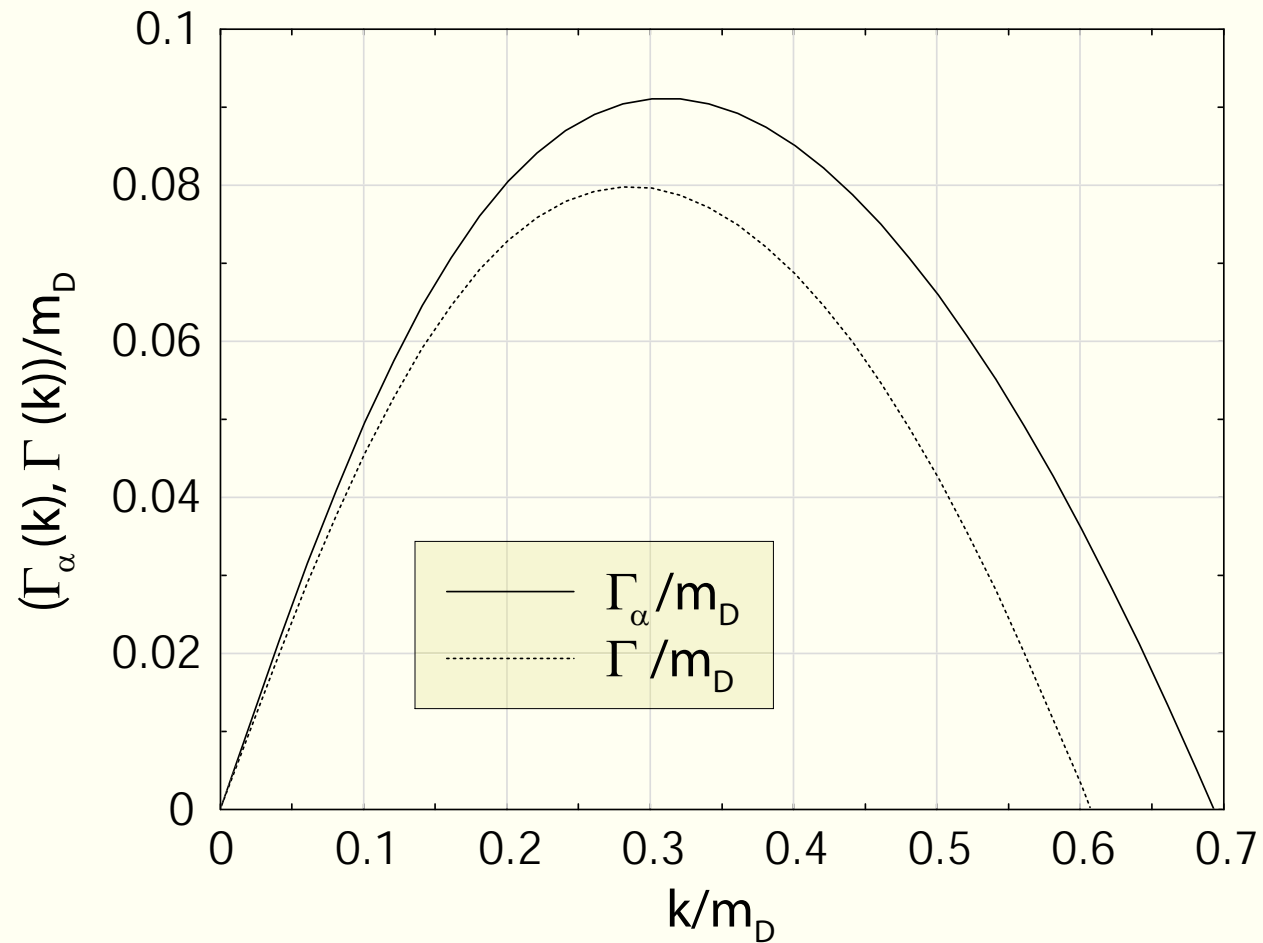
Angular dependence of m_α^2 , m_+^2 , and m_-^2 at fixed $\xi = 10$.

Anisotropic Collective Modes ($\xi > 0$)



Anisotropic poles ($\xi > 0$).

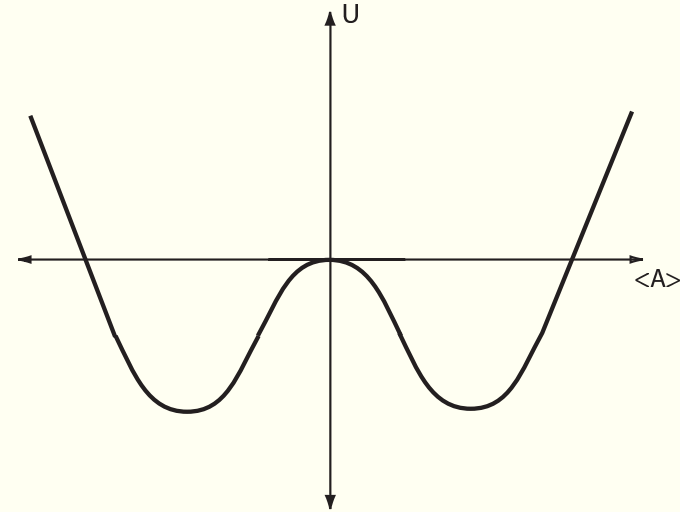
Unstable Modes – $\xi > 0$



$\Gamma_\alpha(k)$ and $\Gamma_-(k)$ as a function of k for $\xi = 10$ and $\theta_n = \pi/8$.

Anisotropic HL Effective Action

Using the requirement of gauge invariance it is possible to determine all n -point functions.



$$\begin{aligned}
 S_{\text{HL}} &= -\frac{g^2}{2} \int_x \int_{\mathbf{p}} f(\mathbf{p}) F_{\mu\nu}^a(x) \left(\frac{p^\nu p^\rho}{(p \cdot D)^2} \right)_{ab} F_\rho^{b\mu}(x) \\
 &= -\frac{g^2}{2} \int_x \int_{\mathbf{p}} f(\mathbf{p}) W^\mu(x, \hat{\mathbf{p}}) W_\mu(x, \hat{\mathbf{p}})
 \end{aligned}$$

For example, from this we can obtain the anisotropic 3-gluon vertex

$$\Gamma^{\mu\nu\lambda}(k, q, r) = \frac{g^2}{2} \int_{\mathbf{p}} \frac{\partial f(\mathbf{p})}{\partial p^\beta} \hat{p}^\mu \hat{p}^\nu \hat{p}^\lambda \left(\frac{r^\beta}{\hat{p} \cdot q \hat{p} \cdot r} - \frac{k^\beta}{\hat{p} \cdot k \hat{p} \cdot q} \right)$$

Real-Time Lattice Simulation

Numerically solve the equations of motion resulting from the HL effective action on a space + velocity lattice.

$$j^\mu[A] = -g^2 \int_{\mathbf{p}} \frac{1}{2|\mathbf{p}|} p^\mu \frac{\partial f(\mathbf{p})}{\partial p^\beta} W^\beta(x; \mathbf{v})$$

with

$$[v \cdot D(A)] W_\beta(x; \mathbf{v}) = F_{\beta\gamma}(A) v^\gamma$$

and $v^\mu = p^\mu / |\mathbf{p}| = (1, \mathbf{v})$.

This has to be solved with

$$D_\mu(A) F^{\mu\nu} = j^\nu$$

where $\nu = 0$ is the Gauss law constraint.

\vec{v} -discretized equations of motion

Recall,

$$j^\nu[A] = -g^2 \int_{\mathbf{p}} \frac{1}{2|\mathbf{p}|} p^\nu \frac{\partial f(\mathbf{p})}{\partial p^\beta} W^\beta(x; \mathbf{v})$$

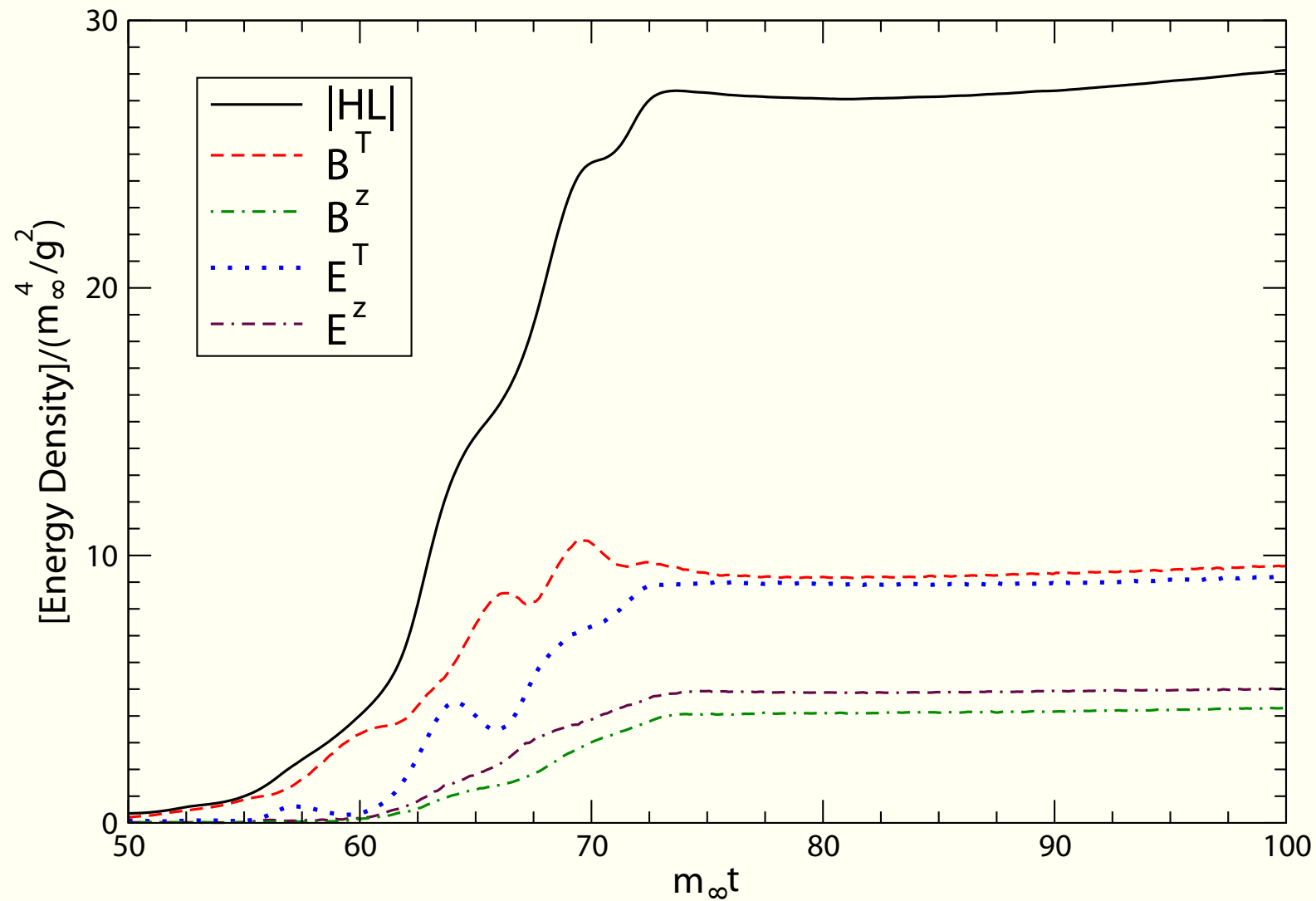
A closed set of gauge-covariant equations is obtained when the angular integral over $\hat{\mathbf{p}}$ is discretized.

The full HL dynamics is then approximated by the following set of equations

$$\begin{aligned} [v \cdot D(A)] \mathcal{W}_{\mathbf{v}} &= (a_{\mathbf{v}} F^{0\mu} + b_{\mathbf{v}} F^{z\mu}) v_\mu \\ D_\mu(A) F^{\mu\nu} &= j^\nu = \frac{1}{\mathcal{N}} \sum_{\mathbf{v}} v^\nu \mathcal{W}_{\mathbf{v}} \end{aligned}$$

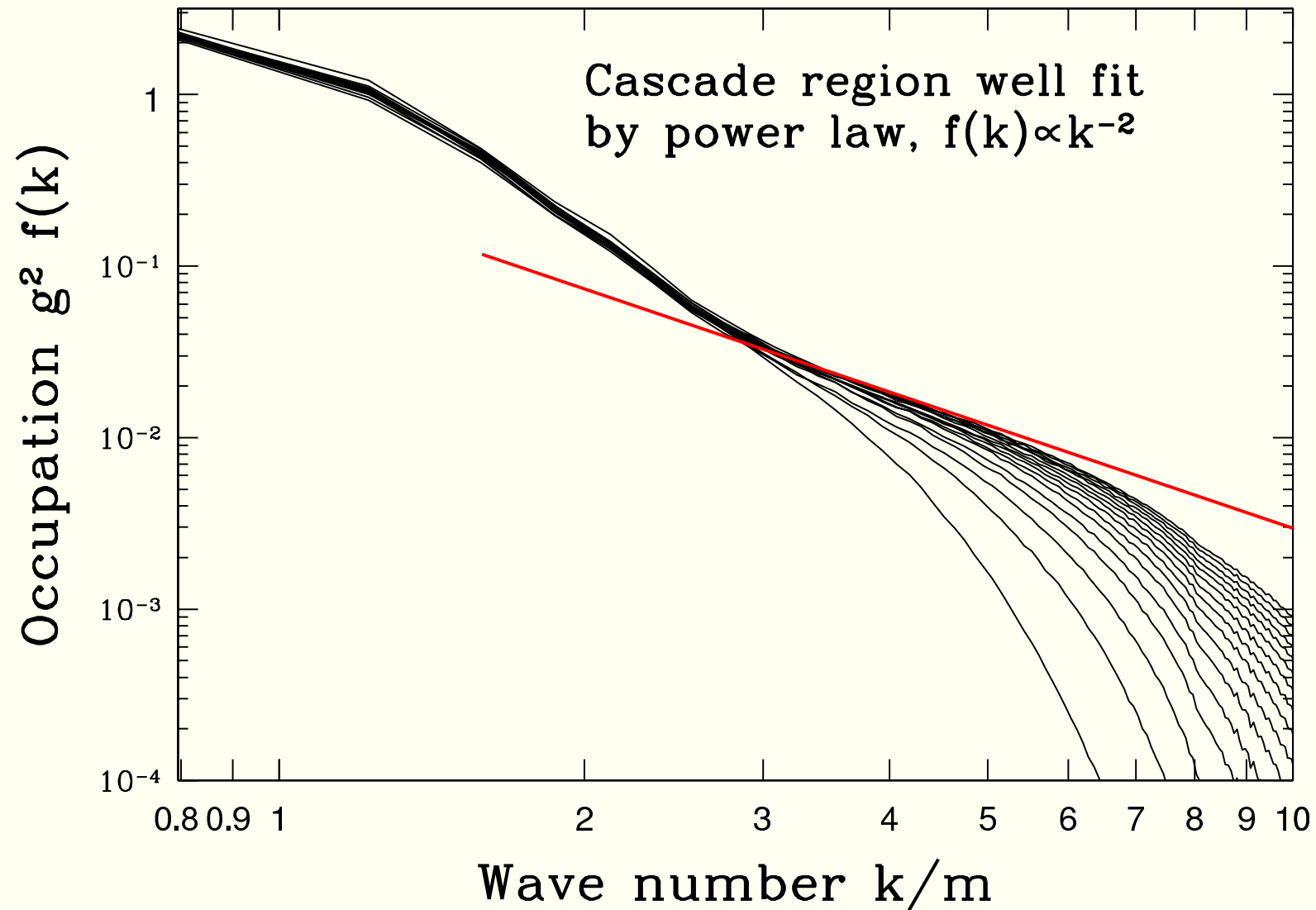
which can be systematically improved by increasing \mathcal{N} .

$3s \times 3v$ Hard-loop results – $\xi = 10$



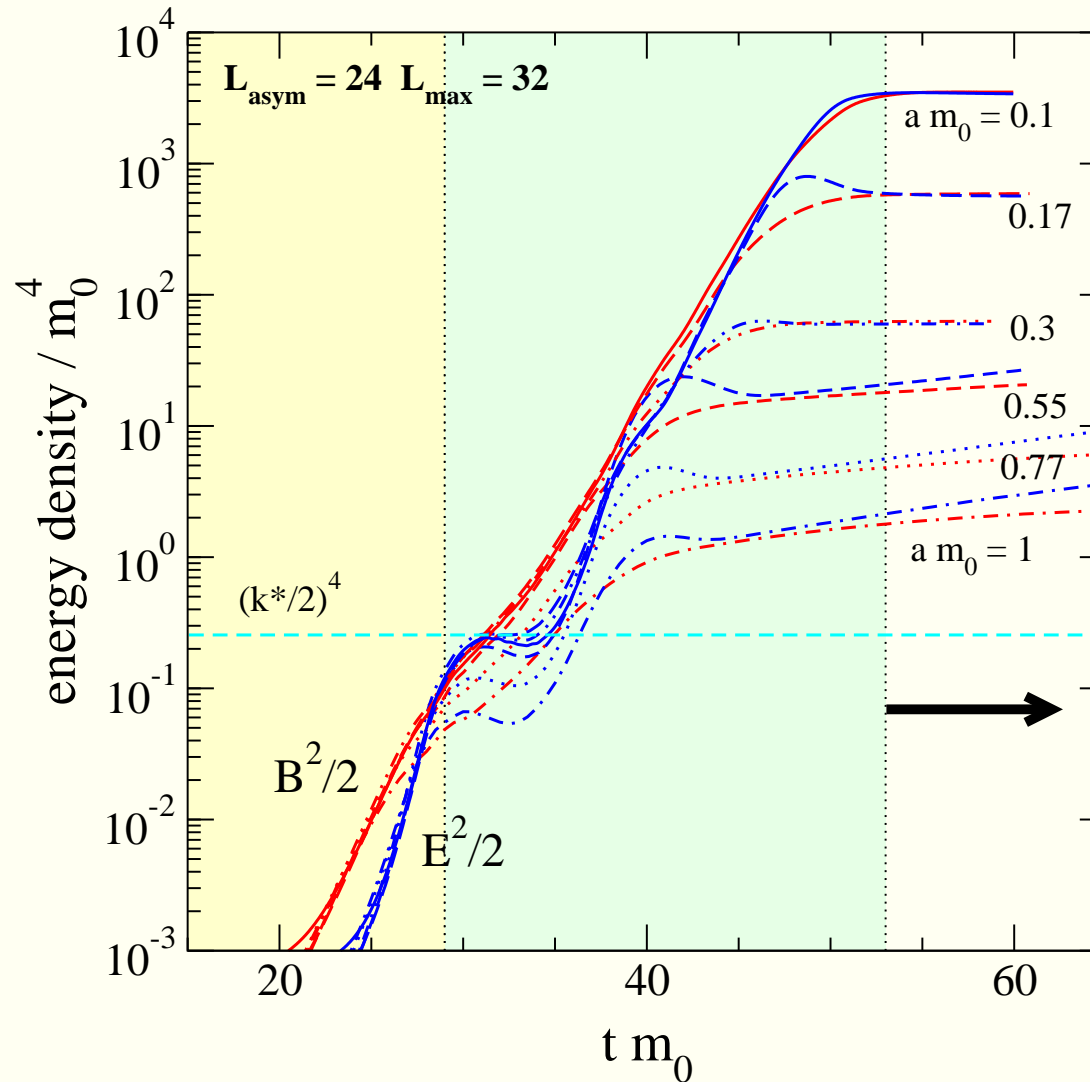
A. Rebhan, P. Romatschke, M. Strickland, hep-ph/0505261

$3s \times 3v$ Hard-loop results – Nonabelian cascade



P. Arnold and G. Moore, hep-ph/0509206; hep-ph/0509226.

3s \times 3v - Larger Anisotropies - $\xi = 100$



Wed May 10 16:09:18 2006

A model of the effect of collisions

We can model the collisional kernel by a Bhatnagar-Gross-Krook (BGK) collision term resulting in a linearized Boltzmann-Vlasov equation of the form

$$[V \cdot D_X, \delta f(p, X)] + g V_\mu F^{\mu\nu} \partial_\nu^{(p)} f(\mathbf{p}) = L(C_{\text{BGK}}[f + \delta f])$$

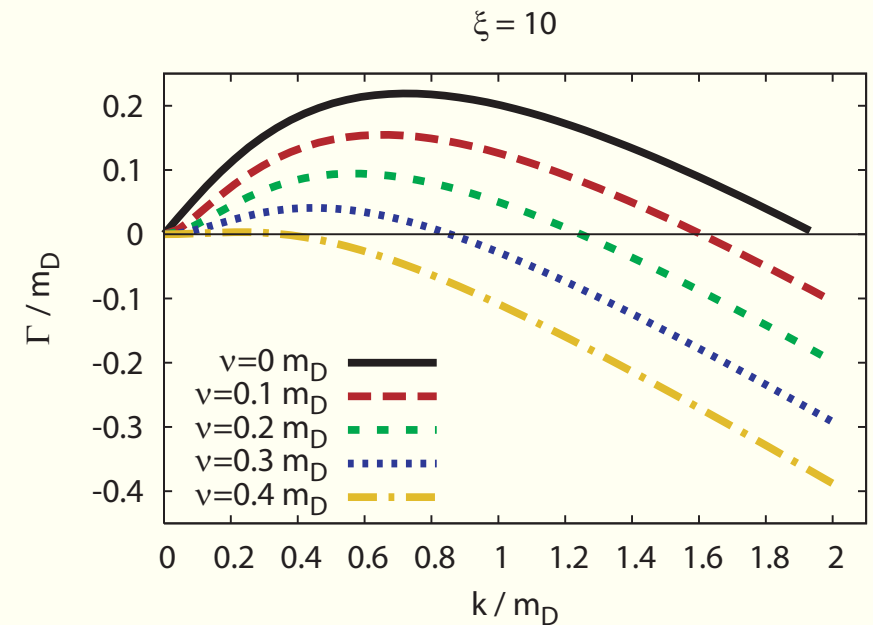
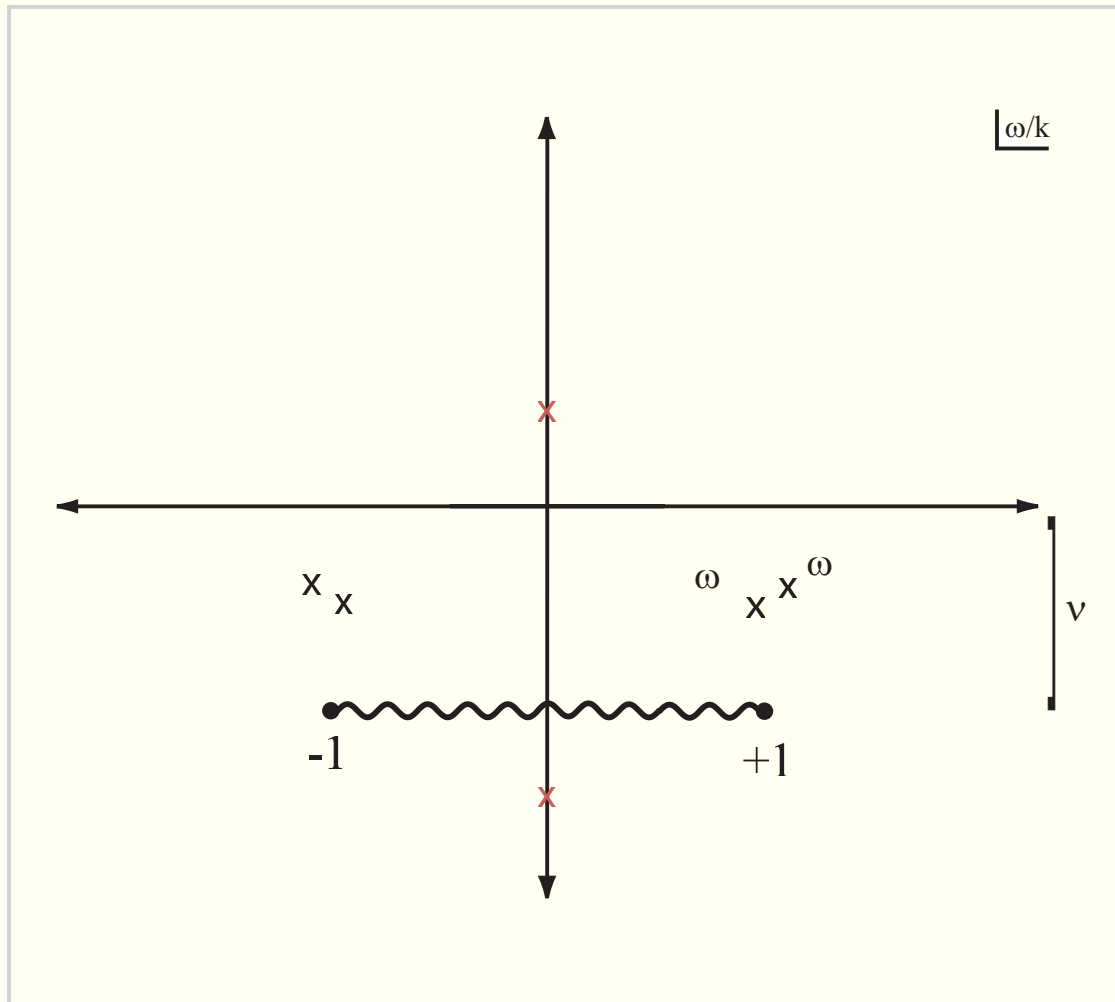
$$C_{\text{BGK}}[f] = -\nu \left(f(p, X) - \frac{N^i(X)}{N_{\text{eq}}^i} f_{\text{eq}}(|\mathbf{p}|) \right)$$

where ν has dimensions of energy and represents the collisional frequency.

For hard, O(1) direction changing interactions $\nu_{\text{hard}}/p_{\text{hard}} \sim g^4 \log g$ and for small-angle ($\theta \sim g$) deflections $\nu_{\text{soft}}/p_{\text{hard}} \sim g^2 \log g$. Assuming the later as an upper bound then $\nu \sim 0.2 m_D$ when $\alpha_s = 0.3$.

B. Schenke, C. Greiner, M. Thoma, and MS, hep-ph/0603029.

BGK Anisotropic Dispersion Relations



In the limit $\xi \rightarrow \infty$ you can show analytically that there is no instability for

$$v > 0.6267 m_D$$

Colored-Particle-in-Cell Simulations (CPIC)

Hard-loop approximation strictly only applies when there is a large scale separation and weak-field limit ($A \ll p_{\text{hard}}/g$).

What happens when one relaxes these assumptions? Let's go back to the transport equations and try to solve without linearization. Recall the Vlasov equation

$$p^\mu [\partial_\mu - g q^a F_{\mu\nu}^a \partial_p^\nu - g f_{abc} A_\mu^b q^c \partial_{q^a}] f(x, p, q) = 0$$

The Vlasov equation is coupled self-consistently to the Yang-Mills equation for the soft gluon fields,

$$D_\mu F^{\mu\nu} = J^\nu = g \int \frac{d^3 p}{(2\pi)^3} dq q v^\nu f(t, \mathbf{x}, \mathbf{p}, q)$$

CPIC - Wong-Yang-Mills equations

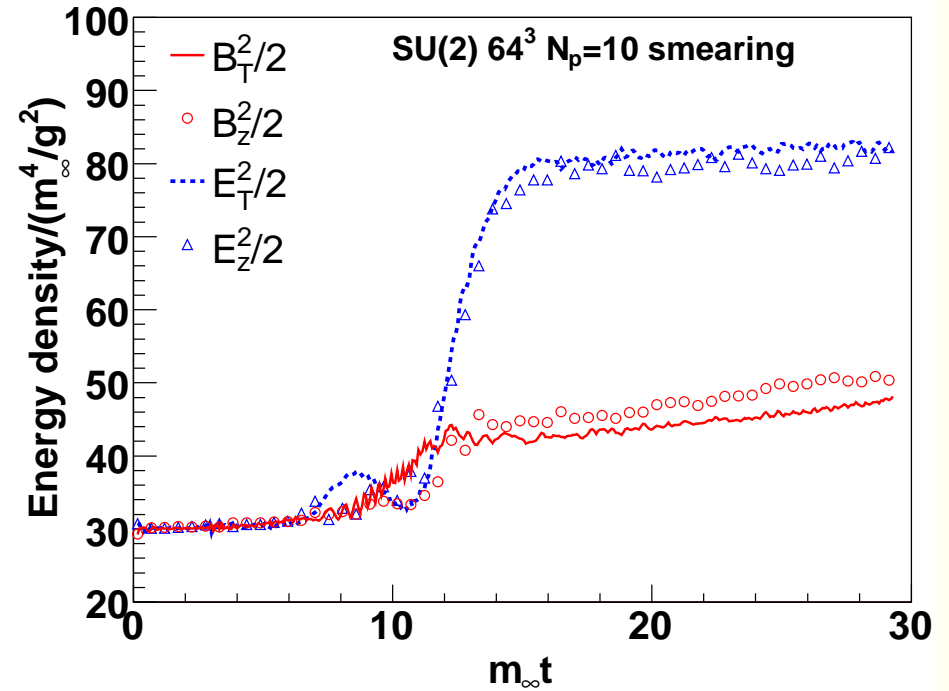
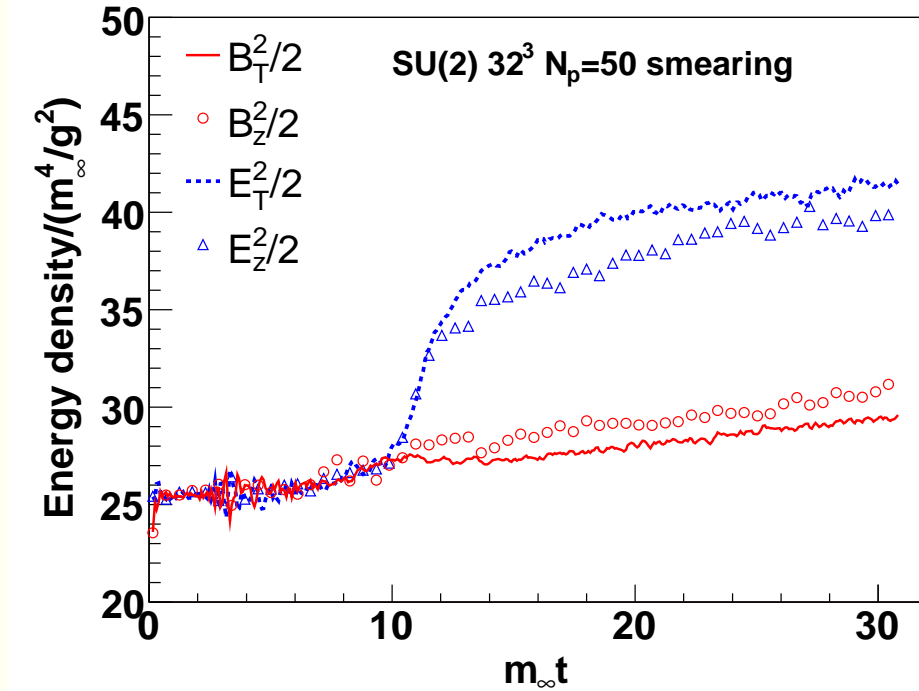
Can be solved numerically by replacing the continuous single-particle distribution $f(\mathbf{x}, \mathbf{p}, q)$ by a large number of test particles:

$$f(\mathbf{x}, \mathbf{p}, q) = \frac{1}{N_{\text{test}}} \sum_i \delta(\mathbf{x} - \mathbf{x}_i(t)) (2\pi)^3 \delta(\mathbf{p} - \mathbf{p}_i(t)) \delta(q^a - q_i^a(t))$$

where $\mathbf{x}_i(t)$, $\mathbf{p}_i(t)$ and $q_i^a(t)$ are the coordinates, momentum, and charge of an individual test particle.

$$\begin{aligned} \frac{d\mathbf{x}_i}{dt} &= \mathbf{v}_i \\ \frac{d\mathbf{p}_i}{dt} &= g q_i^a (\mathbf{E}^a + \mathbf{v}_i \times \mathbf{B}^a) \\ \frac{d\mathbf{q}_i}{dt} &= ig v_i^\mu [A_\mu, \mathbf{q}_i] \\ J^{a\nu} &= \frac{g}{N_{\text{test}}} \sum_i q_i^a v_i^\nu \delta(\mathbf{x} - \mathbf{x}_i(t)) \end{aligned}$$

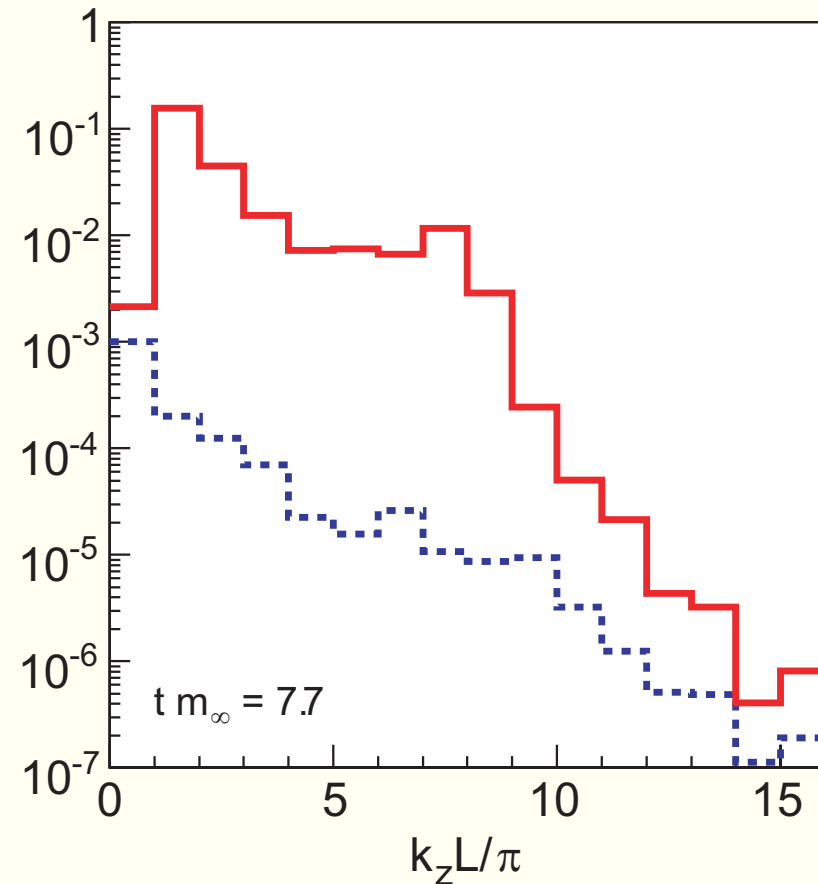
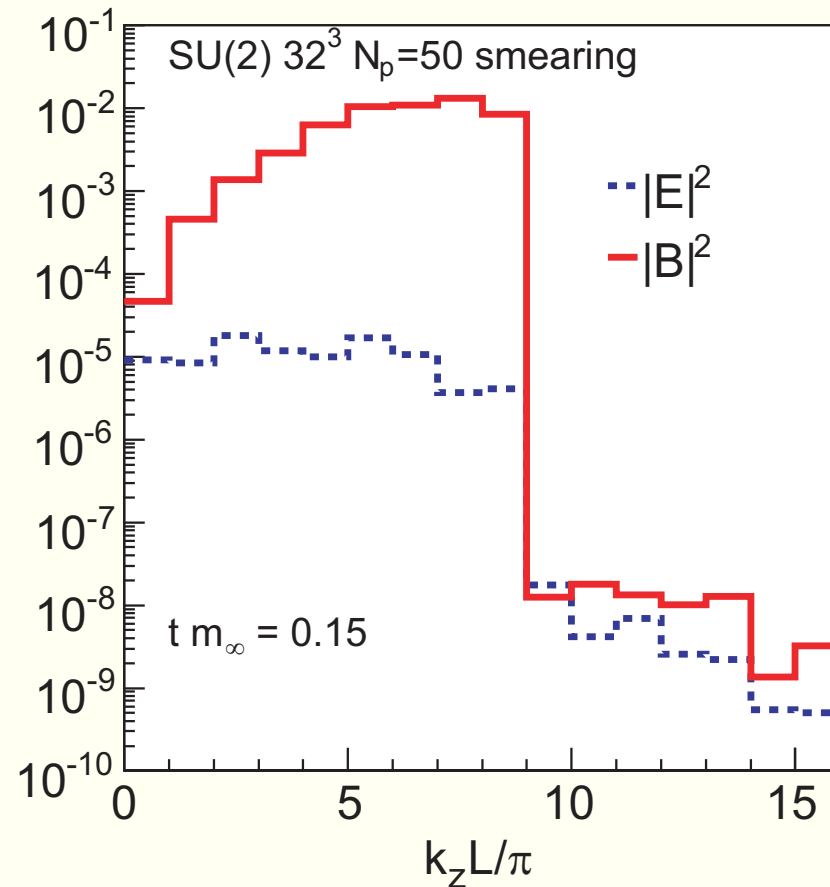
CPIC - Results



Time evolution of the field energy densities for $SU(2)$ gauge group and anisotropic initial particle momentum distributions. Simulation parameters are $L = 5$ fm, $p_{\text{hard}} = 16$ GeV, $g^2 n_g = 10/\text{fm}^3$, $m_\infty = 0.1$ GeV.

A. Dumitru, Y. Nara, and M. Strickland, hep-ph/0604149

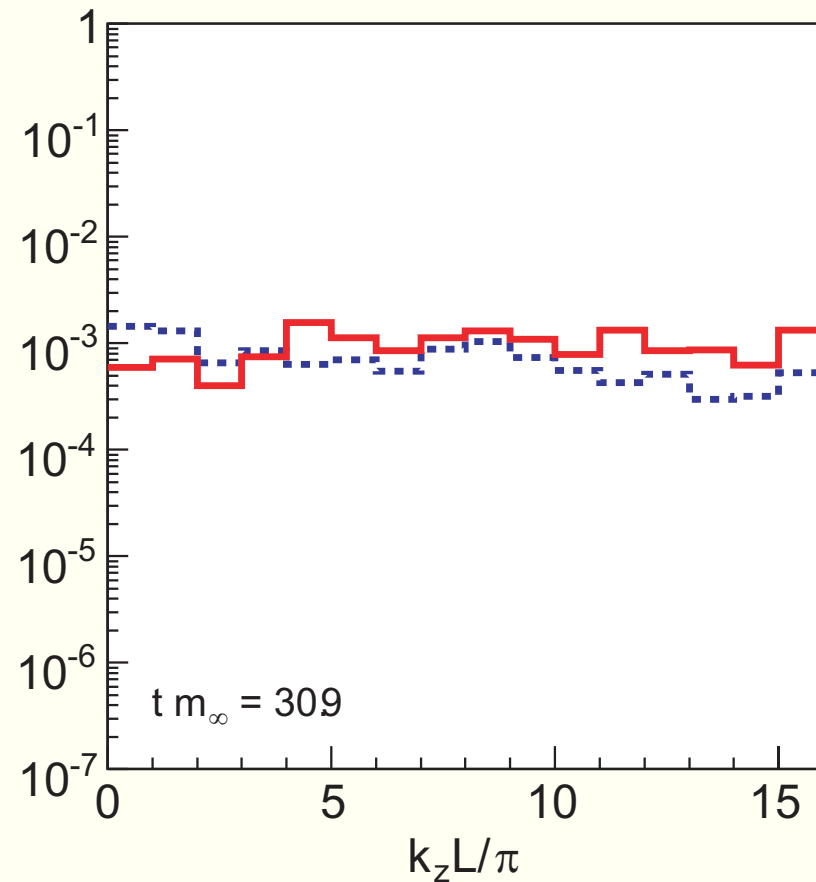
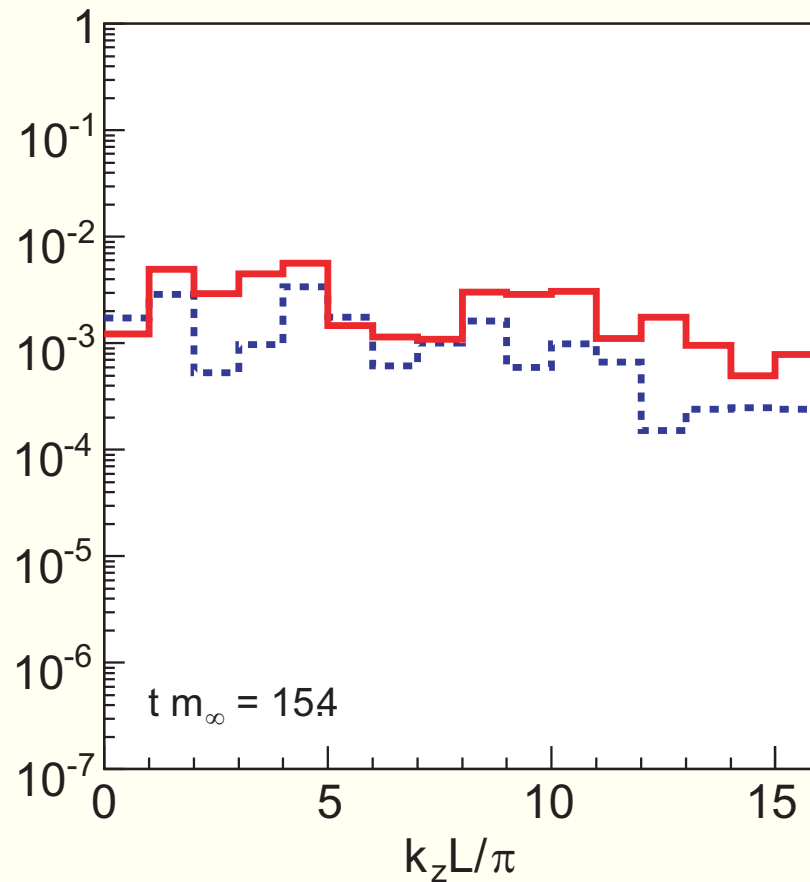
CPIC - Ultraviolet Avalanche



Squares of the Fourier transformed (color-) electric and magnetic fields (in lattice units) at four different times.

A. Dumitru, Y. Nara, and M. Strickland, hep-ph/0604149

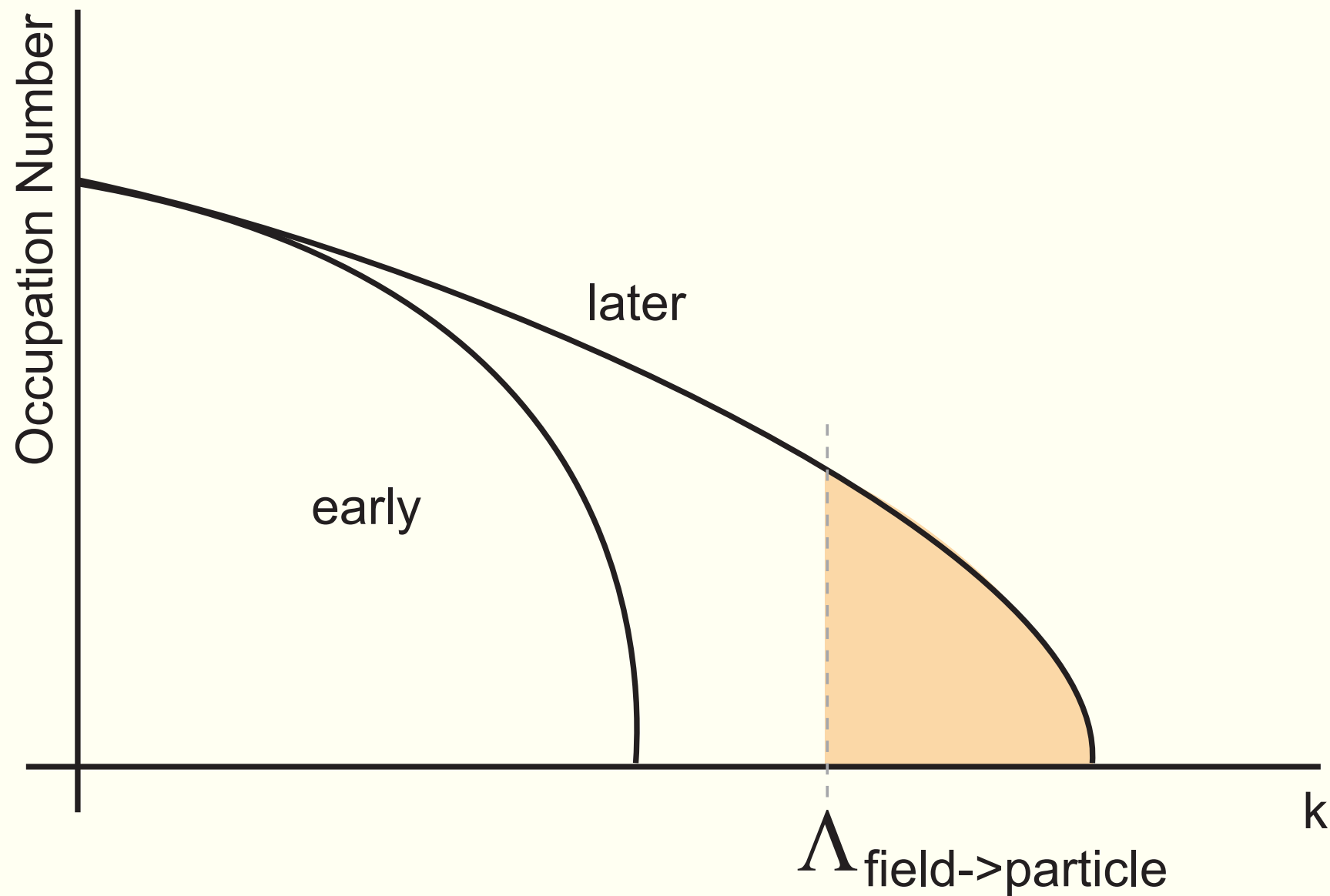
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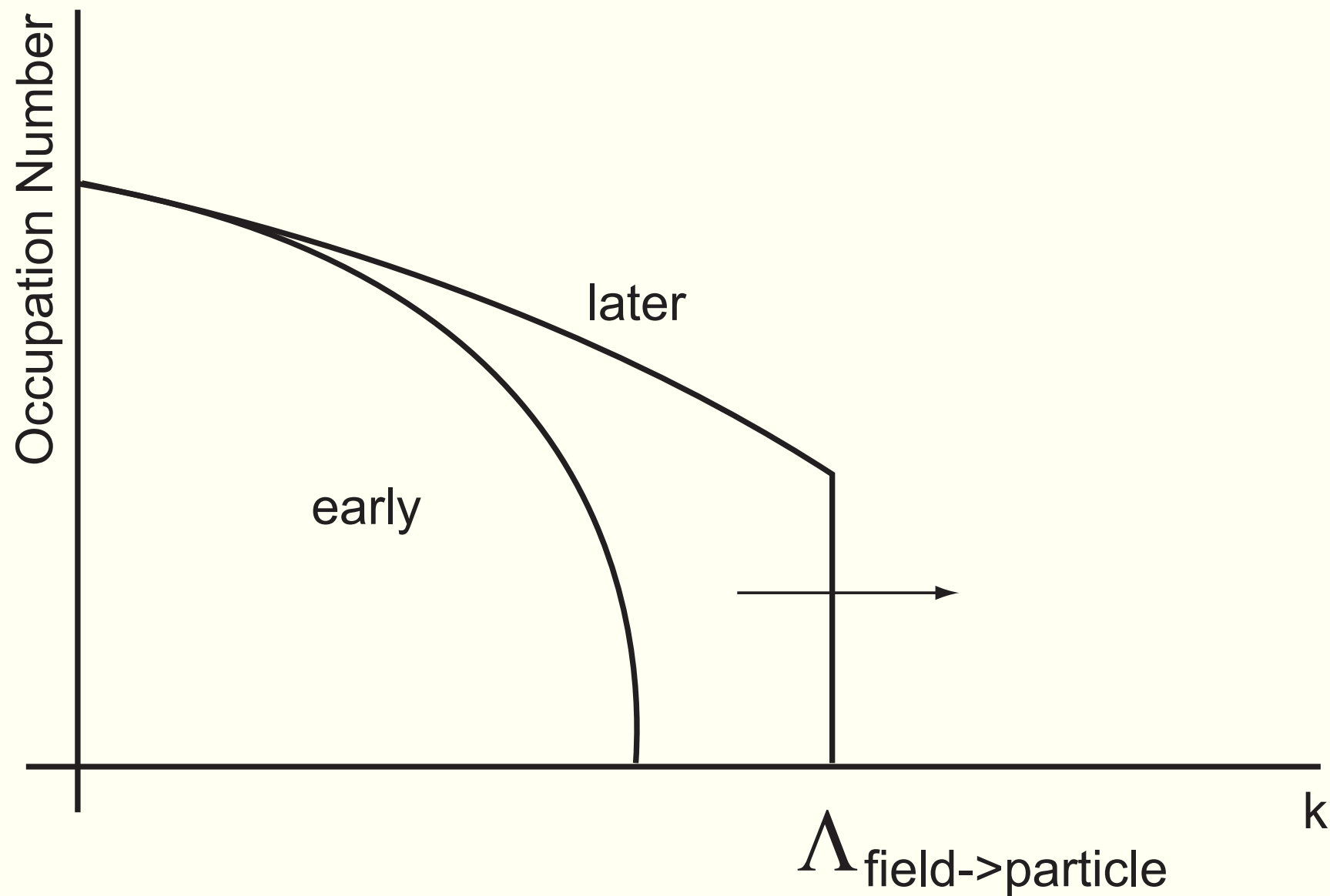
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Cycle of isotropization?



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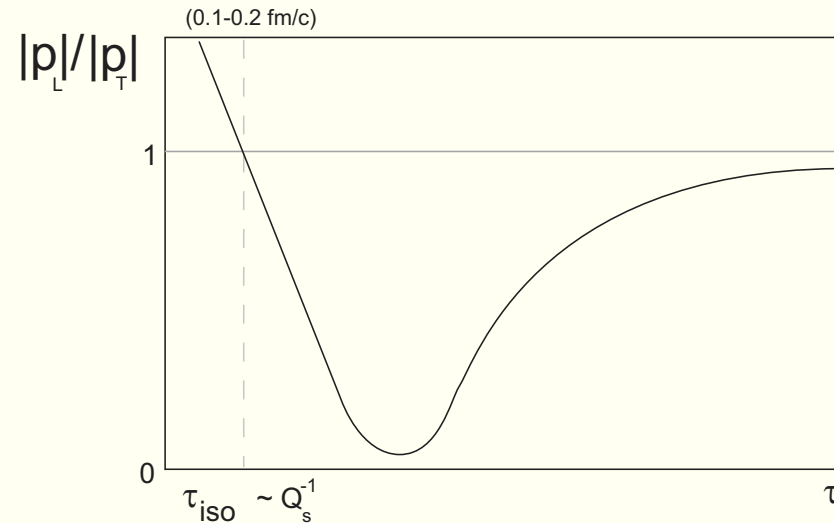
Hard Expanding Loops (HEL)

Including expansion at late times one can show

$$m_{\infty}(\tau) \propto \sqrt{\frac{Q_s}{\tau}}$$

$$A(\tau) \propto e^{\int_{\tau_0}^{\tau} m_{\infty}(t) dt}$$

$$\propto e^{\sqrt{Q_s \tau}}$$



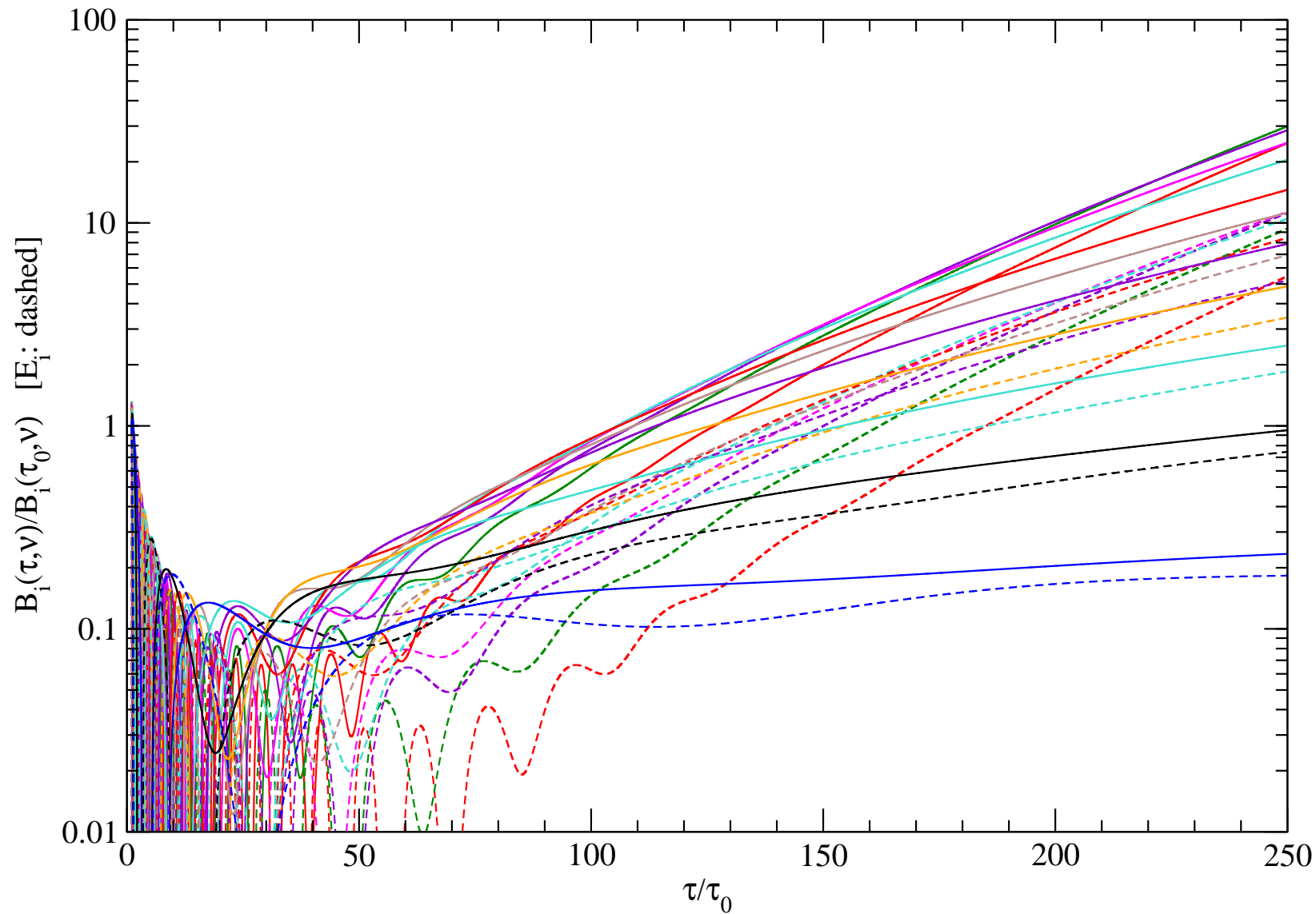
To include the effect of expansion at all times one must solve the Boltzmann-Vlasov equation in an expanding metric using (τ, x^i, η) coordinates (A. Rebhan and P. Romatschke, hep-ph/0605064.):

$$p \cdot D \delta f^a|_{p^\mu} = g p^\alpha F_{\alpha\beta}^a \partial_{(p)}^\beta f_0(\mathbf{p}_\perp, p_\eta)$$

$$\frac{1}{\tau} D_\alpha \left(\tau F^{\alpha\beta} \right) = j^\beta = \frac{g}{2} \int \frac{d^2 p_\perp dy}{(2\pi)^3} p^\beta \delta f$$

Magnetic and electric field strength for each Fourier mode in rapidity

$$c=0.5, \tau_0/\tau_{\text{iso}}=100$$



A. Rebhan and P. Romatschke, hep-ph/0605064.

Other recent related works of interest

- “The Unstable Glasma”, Paul Romatschke, Raju Venugopalan, hep-ph/0605045.
- “Anomalous Viscosity of an Expanding Quark-Gluon Plasma”, M. Asakawa, S.A. Bass, B. Müller, hep-ph/0603092
- Your name here!

Conclusions

- Anisotropic plasmas are qualitatively different than isotropic ones. An entirely new phenomena associated with unstable modes appears.
- For relatively weak anisotropies $3 \text{ space} \times 3 \text{ velocity}$ real-time lattice simulations indicate that for non-abelian plasmas the soft unstable modes “saturate” and the growth then becomes power-law rather than exponential.
- However, for larger anisotropies it appears that exponential field growth can continue similar to an abelian plasma.
- Addition of collisions slows down growth of instabilities but for realistic collisional frequencies instabilities are still present.
- Going beyond the hard-loop approximation by numerically solving the Wong-Yang-Mills equations (CPIC) also shows rapid field growth (but electric fields???) and an “ultraviolet avalanche” accompanied with saturation.